

Newtonian and Post-Newtonian Gravity in Cosmology :

$c \rightarrow \infty$ and Beyond

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Gravity in Cosmology

Einstein gravity (GR) used on large scales and Newtonian gravity used on small scales.

But, Newtonian gravity is incorrect on any scale.

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⇒ We need to go beyond Newtonian theory on small scales

Post-Friedmann Formalism

- Framework for structure formation beyond Newtonian limit
- Uses Post-Newtonian style $\frac{1}{c}$ expansion in Poisson gauge
- PN philosophy in Cosmology different from Solar System:
Metric evolution \Rightarrow Need consistent solution of Einstein Equations
- Linearisation \rightarrow linear perturbation theory (large scales)
- Leading order $\frac{1}{c}$ expansion \rightarrow Newtonian+ regime (small scales)

post-Friedmann metric

- $g_{00} = -\left(1 - \frac{2U_N}{c^2} + \frac{1}{c^4}(2U_N^2 - 4U_P)\right)$
- $g_{0i} = -a\left(\frac{B_i^N}{c^3} + \frac{B_i^P}{c^5}\right)$
- $g_{ij} = a^2\left(\left[1 + \frac{2V_N}{c^2} + \frac{2V_N^2 + 4V_P}{c^4}\right]\delta_{ij} + \frac{h_{ij}}{c^4}\right)$

B_i^N is divergenceless $B_{i,i}^N = 0$

h_{ij} is transverse and traceless $h_{ij,i} = 0, h_{ii} = 0$

At this level, h_{ij} is non-dynamical \rightarrow Not gravitational waves

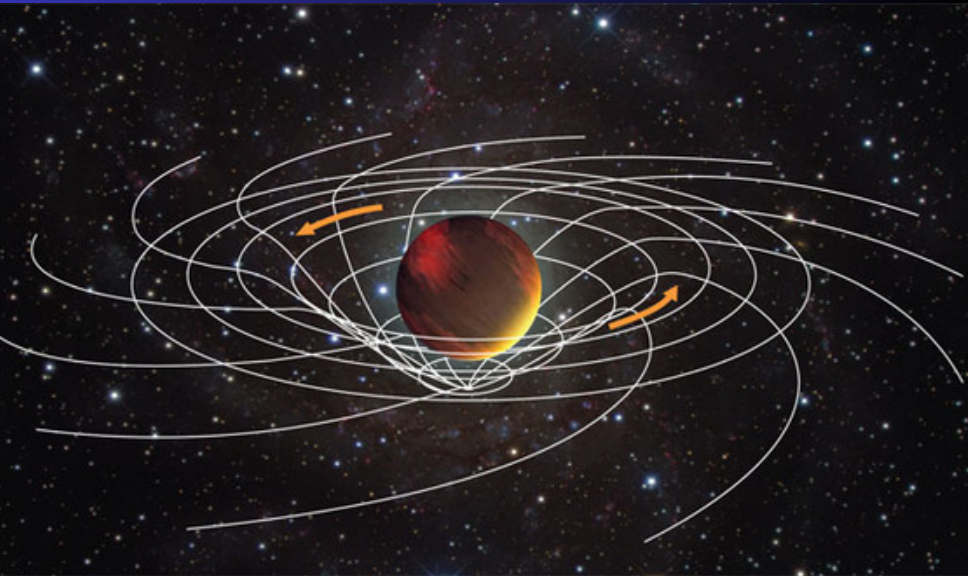
Leading Order Gravitational Equations

$$\frac{1}{c^2} \nabla^2 V_N = -\frac{4\pi G a^2 \rho_b}{c^2} \delta$$

$$\frac{2}{c^2 a^2} \nabla^2 (V_N - U_N) = 0$$

$$\frac{1}{c^3} \nabla^2 B_i^N = -\frac{16\pi G a^2 \rho_b}{c^3} (1 + \delta) v_i + \frac{2}{c^3} \left(\dot{a} U_{N,i} + a \dot{V}_{N,i} \right)$$

Frame Dragging





Frame-Dragging Potential

- What is it physically?

An object's rotation affects spacetime; it “drags” spacetime around

- How can we calculate it¹?

$$\nabla \times \nabla^2 \vec{B}^N = - (16\pi G \rho_b a^2) \nabla \times [(1 + \delta)\vec{v}]$$

Sourced by purely Newtonian quantities

Doesn't affect matter dynamics at this order

Could affect e.g. photon geodesics

¹A similar equation appears in Green&Wald '12, Takada&Futamase '99,

N-body simulations & tessellation

- Theory

An N-body simulation should contain δ and \vec{v}

⇒ We can extract the source term for the Vector potential

- Practice

Extracting velocities non-trivial


Delauney Tessellation Field Estimator (DTFE)²

²Bernardeau, F. & van de Weygaert, R. 1996, MNRAS, 279, 693

Schaap, W. E. & van de Weygaert, R. 2000, A& A, 363, L29 astro-ph/0011007

Schaap, W. PhD thesis "DTFE: the Delaunay Tessellation Field Estimator"

van de Weygaert, R. & Schaap, W. 2009, 665, 291 arXiv: 0708.1441

Cautun, M. & van de Weygaert, R., "The DTFE public software" arXiv: 1105.0370 

Extracted Power Spectra

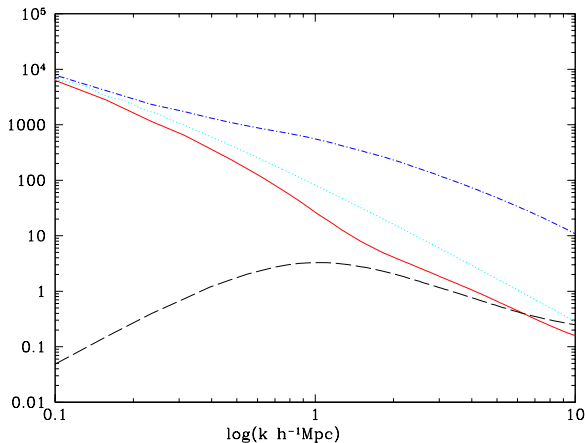


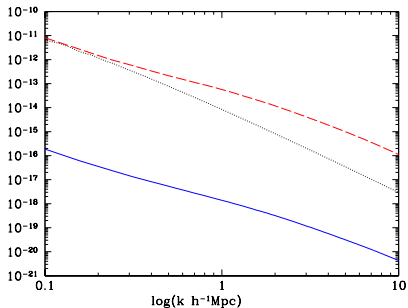
Figure: Power Spectra of density (blue), linear density (cyan), velocity divergence (red) and vorticity (black). Divergence and vorticity divided by $(\mathcal{H}f)^2$

Quiz Time

How much smaller is the power spectrum of B_i^N compared to the Newtonian potential?

- $\sim 10^3$ times smaller
- $\sim 10^5$ times smaller
- $\sim 10^7$ times smaller
- $\sim 10^9$ times smaller

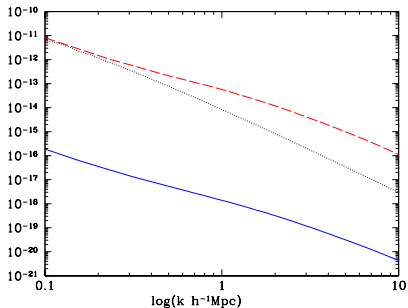
Results



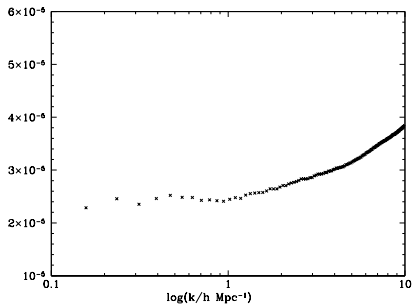
Power spectra of the scalar potential (red)

and vector potential (blue)

Results



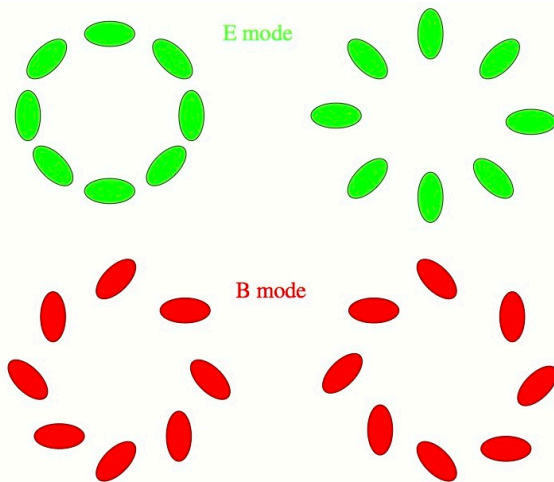
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Ratio of vector potential to
scalar potential

Should we care?

Should we care?



van Waerbeke & Mellier 2003

Conclusions

- We need to go beyond Newtonian gravity on small scales
- Post-Friedmann approach allows calculation of relativistic quantities
- We have computed the vector potential generated in the metric by cosmological large scale structure
 - This is required in the Newtonian regime for consistency of the Einstein equations
 - The first calculation of an intrinsically relativistic quantity in full non-linearity in LSS
- This vector potential could be observed via the weak-lensing B-mode

- Why?

“Normal” estimators-CIC etc, give mass-averaged not volume-averaged velocities

Velocity field artificially set to zero in sparse regions

- How?

Constructs tetrahedra, with nodes located at the particles' positions

No particle inside the circumsphere of another tetrahedron

Velocities interpolated across tetrahedra $\Rightarrow \vec{v}$ known everywhere

Field sampled at random points within grid cell and averaged

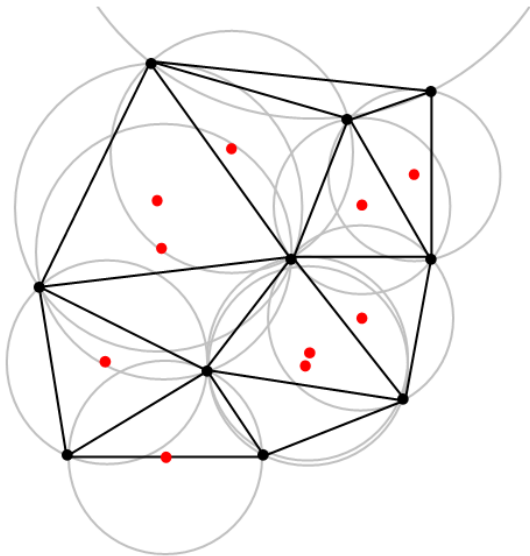


Figure: 2D Delaunay tessellation of black points. Centres of circumcircles in red.